

THREE-DIMENSIONAL SUPERFIELD SUPERGRAVITIES

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We review the superfield formalisms of the three-dimensional supergravities. The $\mathcal{N} = 1, d = 3$ gravitational superfields are generated by the covariant spinor derivative for the tangent $GL(2, R)$ group. The basic gauge group of the $\mathcal{N} = 2, d = 3$ supergravity is defined in the chiral superspace.

The real gravitational $\mathcal{N} = 2$ superfield $h^m(x, \theta, \bar{\theta})$ describes the embedding of the real gravitational superspace into the complex chiral superspace by the analogy with the Ogievetsky-Sokatchev formalism of the $\mathcal{N} = 1, d = 4$ supergravity. The simplest superconformal compensator is the chiral superfield.

The basic gauge group of the $\mathcal{N} = 3, d = 3$ supergravity is defined in the corresponding analytic harmonic superspace. We use the $SU(2)/U(1)$ harmonics by the analogy with the harmonic-superspace formalism of the $\mathcal{N} = 2, d = 4$ supergravity. The harmonic gauge superfields are defined in the decomposition of the covariant harmonic derivative. The superconformal compensator is the $\mathcal{N} = 3$ analytic hypermultiplet superfield.

4-DIMENSIONAL SUPERFIELD SUPERGRAVITIES

$\mathcal{N} = 1, d = 4$ superfield supergravity

Wess, Zumino; Ogievetsky, Sokatchev; Gates, Grisaru, Roček, Siegel

Chirality preserving constraints for the supervielbein matrix E_M^A were solved in terms of the $\mathcal{N} = 1$ axial-vector gauge superfield $H^m(x, \theta, \bar{\theta})$. The nonlinear supergravity superdeterminant action $\kappa^{-2} \int d^4x d^2\theta d^2\bar{\theta} E(H^m)$ was constructed and quantized.

$\mathcal{N} = 2, d = 4$ superfield supergravity in the harmonic $SU(2)/U(1)$ superspace was constructed by Galperin, Ivanov, Kalitzin, Ogievetsky, Sokatchev

The gauge supergravity superfields and the $\mathcal{N} = 2, d = 4$ matter superfields live in the Grassmann-analytic superspace. The supergravity-matter classical superfield actions were constructed in this approach, but the quantum superfield calculations were not developed in this formalism.

We review the superfield formalisms of the simplest three-dimensional supergravities.

FLAT 3-DIMENSIONAL SUPERSPACES

$\mathcal{N} = 1$ superspace: $z = (x^m, \theta^\mu)$, $m = 0, 1, 2$, $\mu = 1, 2$

Spinor derivatives

$$D_\mu = \partial_\mu + i\theta^\nu(\gamma^m)_{\mu\nu}\partial_m, \quad \partial_\mu\theta^\nu = \delta_\mu^\nu, \quad \partial_m x^n = \delta_m^n$$

$$\gamma^m\gamma^n = -\eta^{mn}I + \varepsilon^{mnp}\gamma_p, \quad \eta^{mn} = \mathbf{diag}(1, -1, -1)$$

$$\delta_\epsilon x^m = -i(\epsilon\gamma^m\theta), \quad \delta_\epsilon\theta^\beta = \epsilon^\beta$$

$\mathcal{N} = 1$ scalar superfield : $\phi(z)$

$\mathcal{N} = 1$ Maxwell superfield : $A_\mu(z)$, $\delta_\lambda A_\mu = D_\mu\lambda(z)$

$\mathcal{N} = 2, d = 3$ superspace is analogous to the $\mathcal{N} = 1, d = 4$ superspace:

$z = (x^m, \theta^\mu, \bar{\theta}^\mu)$, $m = 0, 1, 2$, $\mu = 1, 2$

Spinor derivatives

$$D_\mu = \partial_\mu + i\bar{\theta}^\nu(\gamma^m)_{\mu\nu}\partial_m, \quad \bar{D}_\mu = -\bar{\partial}_\mu - i\theta^\nu(\gamma^m)_{\mu\nu}\partial_m$$

Chiral $\mathcal{N} = 2$ superspace: $\zeta = (x_L^m, \theta^\mu)$, $x_L^m = x^m + i(\theta\gamma^m\bar{\theta})$

$\mathcal{N} = 2$ scalar chiral superfield : $\phi(x_L^m, \theta^\mu)$

Anti-chiral superspace: $\bar{\zeta} = (x_R^m, \bar{\theta}^\mu)$, $x_R^m = x^m - i(\theta\gamma^m\bar{\theta})$

$\mathcal{N} = 2, d = 3$ abelian gauge superfield:

$$\delta_\lambda V(x, \theta, \bar{\theta}) = i\lambda(\zeta) - i\bar{\lambda}(\bar{\zeta}), \quad W = D^\alpha \bar{D}_\alpha V$$

$\mathcal{N} = 3, d = 3$ superspace: $z = (x^m, \theta_{(kl)}^\mu), \quad k, l = 1, 2$

Automorphism group is $SU(2)$ and we can use the $SU(2)/U(1)$ harmonics u_k^\pm and the harmonic derivatives

$$\partial^{++} u_k^- = u_k^+, \quad \partial^{--} u_k^+ = u_k^-, \quad \partial^0 u_k^\pm = \pm u_k^\pm$$

Harmonic projections of spinor coordinates

$$\theta^{\mu\pm\pm} = \theta_{(kl)}^\mu u^{k\pm} u^{l\pm}, \quad \theta^{\mu 0} = \theta_{(kl)}^\mu u^{k+} u^{l-}$$

Analytic $\mathcal{N} = 3$ superspace: $\zeta = (x_A^m, \theta^{\mu++}, \theta^{\mu 0}, u)$

Analytic hypermultiplet: $q^+(\zeta)$

Maxwell $\mathcal{N} = 3$ superfield: $\delta_\lambda V^{++}(\zeta) = D^{++} \lambda(\zeta)$

$\mathcal{N} = 1, d = 3$ SUPERGRAVITY
 Gates, Grisaru, Roček, Siegel; Zupnik, Pak

Superdiffeomorphism group

$$\delta_\xi x^m = \xi^m(z), \quad \delta_\xi \theta^\mu = \xi^\mu(z)$$

Holonomic basis $\partial_M = (\partial_m, \partial_\mu), \quad \delta_\xi \partial_M = -\partial_M \xi^N \partial_N$

Flat-superspace basis $D_M = (\partial_m, D_\mu)$

The $\mathcal{N} = 1$ supergravity spinor differential operator contains the superconformal gauge superfields

$$\Delta_\alpha = D_\alpha + i h_\alpha^m(z) \partial_m + h_\alpha^\mu(z) D_\mu, \quad \delta \Delta_\alpha = -\frac{1}{2} \sigma(z) \Delta_\alpha - \lambda_\alpha^\beta(z) \Delta_\beta$$

where $\sigma(z)$ is the superconformal parameter and $\lambda_\alpha^\beta(z)$ describe the $SL(2, R)$ gauge transformations.

Superconformal gauge condition:

$$h_\alpha^\mu(z) = 0, \quad \Delta_\alpha = D_\alpha + i h_\alpha^m \partial_m$$

The composed $SL(2, R)$ and Weyl parameters are induced by the superdiffeomorphism parameters

$$\tilde{\sigma} = \Delta_\alpha \xi_0^\alpha, \quad \tilde{\lambda}_\alpha^\mu = \Delta_\alpha \xi_0^\mu - \frac{1}{2} \delta_\alpha^\mu \Delta_\beta \xi_0^\beta,$$

$$\delta \Delta_\alpha = -\Delta_\alpha \xi_0^\rho \Delta_\rho = -(D_\alpha + i h_\alpha^n \partial_n) \xi_0^\rho \Delta_\rho$$

The gauge transformations in this gauge are nonlinear in the basic superfields

$$\delta h_\alpha^m = -i D_\alpha \xi_0^m - 2(\gamma^m)_{\alpha\beta} \xi_0^\beta + h_\alpha^n \partial_n \xi_0^m - [(D_\alpha + i h_\alpha^n \partial_n) \xi_0^\beta] h_\beta^m$$

where $\xi_0^m = \xi^m - i \xi^\mu \theta^\nu (\gamma^m)_{\mu\nu}, \quad \xi_0^\mu = \xi^\mu$

Scalar compensator of the $\mathcal{N} = 1$ Poincaré supergravity

$$\Sigma(z) = 1 + \kappa \Phi(z), \quad \kappa \text{ is the gravitational constant}$$

$$\delta \Phi(z) = \frac{1}{2} \Delta_\alpha \xi_0^\alpha [\kappa^{-1} + \Phi(z)],$$

$$\mathcal{D}_\alpha = \Sigma \Delta_\alpha = G_\alpha^M D_M, \quad \delta \mathcal{D}_\alpha = -\tilde{\lambda}_\alpha^\beta \mathcal{D}_\beta$$

We introduce the $SL(2, R)$ spinor connection $\Omega_{\alpha, \beta}^\rho$

$$\delta \Omega_{\alpha, \beta}^\rho = -\mathcal{D}_\alpha \tilde{\lambda}_\alpha^\rho + (\lambda \Omega_{\alpha, \beta}^\rho)$$

The covariant vector covariant operator is

$$\mathcal{D}_a = -\frac{i}{4} [\{\mathcal{D}_\alpha, \mathcal{D}_\beta\} - (\Omega_{\alpha, \beta}^\rho + \Omega_{\beta, \alpha}^\rho) \mathcal{D}_\rho] = G_a^M D_M$$

where G_a^M, G_α^M are the composed supervielbein matrix.

The inverse supervielbein matrix satisfies the relations

$$E_M^A G_B^M = \delta_B^A, \quad E = \mathbf{Ber} E_M^A$$

$$\delta E = -(\partial_m \xi^m - \partial_\mu \xi^\mu) E, \quad \delta(d^5 z E) = 0$$

The composed vector superfield connection has the form

$$\Gamma_{a, \pi}^\rho = -\frac{i}{2} (\gamma_a)^{\alpha\beta} \{ \Delta_\alpha \Omega_{\beta, \pi}^\rho - \Omega_{\alpha, \beta}^\varphi \Omega_{\varphi, \pi}^\rho - \Omega_{\alpha, \pi}^\varphi \Omega_{\beta, \varphi}^\rho \}$$

The corresponding covariant derivative is

$$\nabla_a \mathcal{D}_\pi = \mathcal{D}_a \mathcal{D}_\pi - \Gamma_{a, \pi}^\rho \mathcal{D}_\rho, \quad \nabla_\pi \mathcal{D}_a = \mathcal{D}_\pi \mathcal{D}_a - (\gamma_a)_{\rho\xi} (\gamma^b)^{\varphi\rho} \Omega_{\pi, \varphi}^\rho \mathcal{D}_b$$

The covariant supergravity constraints have the form

$$\nabla_\alpha \mathcal{D}_\beta + \nabla_\beta \mathcal{D}_\alpha = 2i (\gamma^a)_{\alpha\beta} \mathcal{D}_a$$

$$\nabla_a \mathcal{D}_\pi - \nabla_\pi \mathcal{D}_a = R (\gamma_a)_\pi^\rho \mathcal{D}_\rho,$$

where $R(z)$ is the basic scalar superfield.

The simplest $\mathcal{N} = 1$ supergravity action is

$$\frac{1}{\kappa} \int dz^5 E(z) R(z)$$

$\mathcal{N} = 2, d = 3$ SUPERGRAVITY

$\mathcal{N} = 2$ supergravity formalism is based on the three-dimensional version of the Ogievetsky-Sokatchev approach. We consider the superdiffeomorphism group in the $\mathcal{N} = 2, d = 3$ chiral superspace

$$\delta x_L^m = \lambda^m(\zeta), \quad \delta \theta^\mu = \lambda^\mu(\zeta), \quad \delta \bar{\partial}_\mu = -(\bar{\partial}_\mu \bar{\lambda}^\nu) \bar{\partial}_\nu$$

This conformal $\mathcal{N} = 2$ supergravity transformations preserve chirality.

In the anti-chiral basis we have analogous conjugated relations

$$\delta x_R^m = \bar{\lambda}^m(\bar{\zeta}), \quad \delta \bar{\theta}^\mu = \bar{\lambda}^\mu(\bar{\zeta}), \quad \delta \partial_\mu = -(\partial_\mu \lambda^\nu) \partial_\nu$$

The real supergravity superspace $Z^M = (x^m, \theta^\mu, \bar{\theta}^\mu)$ is embedded into the chiral superspace

$$x_L^m = x^m + iH^m(x, \theta, \bar{\theta}), \quad x_R^m = x^m - iH^m(x, \theta, \bar{\theta})$$

where $H^m(x, \theta, \bar{\theta}) = (\theta \gamma^m \bar{\theta}) + h^m(x, \theta, \bar{\theta})$ is the gravitational axial-vector superfield.

The passive gauge transformations of the $\mathcal{N} = 2$ conformal supergravity have the form

$$\begin{aligned} \delta x^m &= \frac{1}{2} \lambda^m(x + iH, \theta) + \frac{1}{2} \bar{\lambda}^m(x - iH, \bar{\theta}) \\ \delta \theta^\mu &= \lambda^\mu(x + iH, \theta), \quad \delta \bar{\theta}^\mu = \bar{\lambda}^\mu(x - iH, \bar{\theta}) \\ \delta H^m &= \frac{i}{2} \bar{\lambda}^m(x - iH, \bar{\theta}) - \frac{i}{2} \lambda^m(x + iH, \theta) \end{aligned}$$

We consider the flat chiral basis

$$\begin{aligned} x_{0L}^m &= x^m + i(\theta \gamma^m \bar{\theta}), \quad \zeta_0 = (x_{0L}^m, \theta^\mu) \\ \lambda^m(x + iH, \theta) &= T(ih) \lambda^m(\zeta_0), \quad \lambda^\mu(x + iH, \theta) = T(ih) \lambda^\mu(\zeta_0), \\ T(ih) &= 1 + ih^m \partial_m - \frac{1}{2} h^m h^n \partial_m \partial_n - \frac{i}{6} h^m h^n h^r \partial_m \partial_n \partial_r + \dots \end{aligned}$$

The covariant spinor derivatives in the central coordinates $x^m, \theta^\mu, \bar{\theta}^\mu$ have the form

$$\begin{aligned}\Delta_\alpha &= D_\alpha + i\Delta_\alpha h^m \partial_m = D_\alpha + iD_\alpha h^n [(I - i\partial h)^{-1}]_n^m \partial_m, \\ \bar{\Delta}_\alpha &= \bar{D}_\alpha - i\bar{\Delta}_\alpha h^m \partial_m = \bar{D}_\alpha - i\bar{D}_\alpha h^n [(I + i\partial h)^{-1}]_n^m \partial_m\end{aligned}$$

The tangent $GL(2, C)$ transformations of these derivatives are induced by the chiral diffeomorphism transformations

$$\delta\Delta_\alpha = -(\Delta_\alpha \lambda^\beta) \Delta_\beta, \quad \delta\bar{\Delta}_\alpha = (\bar{\Delta}_\alpha \bar{\lambda}^\beta) \bar{\Delta}_\beta$$

The basic superfield blocks $c_{\mu\nu}^m \partial_m$ can be constructed via the anticommutator of spinor covariant derivatives

$$\frac{1}{2}\{\Delta_\mu, \bar{\Delta}_\nu\} = -i\{(\gamma^m)_{\mu\nu} + \frac{1}{2}[\Delta_\mu, \bar{\Delta}_\nu]h^m\}\partial_m = c_{\mu\nu}^m \partial_m$$

We consider the superconformal chiral parameter

$$j = \partial_m^L \lambda^m - \partial_\mu \lambda^\mu, \quad \delta d^3 x_L d^2 \theta = j d^3 x_L d^2 \theta = j d^5 \zeta$$

and introduce the chiral superfield compensator

$$\delta\Phi = -\frac{1}{2}j\Phi, \quad \delta(d^5 \zeta \Phi^2) = 0$$

This chiral compensator allow us to construct the $\mathcal{N} = 2$ Poincaré supergravity action.

$\mathcal{N} = 3, d = 3$ SUPERGRAVITY

We use the analogy with the Galperin-Ivanov-Ogievetsky-Sokatchev formalism in $\mathcal{N} = 2, d = 4$ supergravity. Let us consider the arbitrary transformations of the $\mathcal{N} = 3, d = 3$ harmonic analytic superspace

$$\begin{aligned}\delta x_A^m &= \lambda^m(\zeta), & \delta \theta^{++\mu} &= \lambda^{++\mu}(\zeta), & \delta \theta^{0\mu} &= \lambda^{0\mu}(\zeta), \\ \delta u_k^+ &= \lambda^{++}(\zeta) u_k^-, & \zeta &= (x_A^m, \theta^{++\mu}, \theta^{0\mu}, u_k^\pm),\end{aligned}$$

where $\lambda^m(\zeta), \lambda^{++\mu}(\zeta), \lambda^{0\mu}(\zeta), \lambda^{++}(\zeta)$ are the analytic parameters of the conformal supergravity.

The harmonic and spinor derivatives in the flat analytic superspace are

$$\begin{aligned}\mathcal{D}^{++} &= \partial^{++} + 2i\theta^{++\alpha}\theta^{0\beta}\partial_{\alpha\beta}^A + \theta^{++\alpha}\partial_\alpha^0 + 2\theta^{0\alpha}\partial_\alpha^{++}, & \partial_{\alpha\beta}^A &= (\gamma^m)_{\alpha\beta}\partial_m^A, \\ \mathcal{D}^{--} &= \partial^{--} - 2i\theta^{--\alpha}\theta^{0\beta}\partial_{\alpha\beta}^A + \theta^{--\alpha}\partial_\alpha^0 + 2\theta^{0\alpha}\partial_\alpha^{--}, \\ \mathcal{D}^0 &= \partial^0 + 2\theta^{++\alpha}\partial_\alpha^{--} - 2\theta^{--\alpha}\partial_\alpha^{++}, & [\mathcal{D}^{++}, \mathcal{D}^{--}] &= \mathcal{D}^0, \\ D_\alpha^{++} &= \partial_\alpha^{++}, & D_\alpha^{--} &= \partial_\alpha^{--} + 2i\theta^{--\beta}\partial_{\alpha\beta}^A, & D_\alpha^0 &= -\frac{1}{2}\partial_\alpha^0 + i\theta^{0\beta}\partial_{\alpha\beta}^A, \\ \partial_m^A x_A^n &= \delta_m^n, & \partial_\alpha^0 \theta^{0\beta} &= \delta_\alpha^\beta, & \partial_\alpha^{\mp\mp} \theta^{\pm\pm\beta} &= \delta_\alpha^\beta.\end{aligned}$$

The analytic integral measure is

$$\begin{aligned}d\zeta^{-4} &= \frac{1}{16}d^3x_A(\partial^{--\alpha}\partial_\alpha^{--})(\partial^{0\alpha}\partial_\alpha^0)du, & (0.1) \\ \delta d\zeta^{-4} &= (\partial_m^A \lambda^m + \partial^{--}\lambda^{++} - \partial_\mu^{--}\lambda^{++\mu} - \partial_\mu^0\lambda^{0\mu})d\zeta^{-4} = -2\Lambda d\zeta^{-4}.\end{aligned}$$

Nonanalytic transformations have the form

$$\delta \theta^{--\mu} = \Lambda^{--\mu}(\zeta, \theta^{--}).$$

We define the Killing operator

$$\begin{aligned} K &= \lambda^m \partial_m + \lambda^{++} \partial^{--} + \lambda^{++\mu} \partial_\mu^{--} + \lambda^{0\mu} \partial_\mu^0 + \Lambda^{--\mu} \partial_\mu^{++} \\ &= \Lambda^m \partial_m + \lambda^{++} \mathcal{D}^{--} + \Lambda^{++\mu} D_\mu^{--} + \Lambda^{0\mu} D_\mu^0 + \Lambda^{--\mu} D_\mu^{++} \end{aligned}$$

Constraints for the flat parameters:

$$\begin{aligned} D_\mu^{++} \Lambda^m &= -4i \lambda^{++} (\gamma^m)_{\mu\nu} \theta^{0\nu} + 2i \lambda^{++\nu} (\gamma^m)_{\mu\nu} = 2i (\gamma^m)_{\mu\nu} \Lambda^{++\nu}, \\ D_\nu^{++} \Lambda^{0\mu} &= -2\delta_\nu^\mu \lambda^{++}. \end{aligned}$$

The basic operator of the $\mathcal{N} = 3$ conformal supergravity contains the gauge gravitational superfields

$$\begin{aligned} \Delta^{++} &= \mathcal{D}^{++} + G^{++} + h^\mu D_\mu^{++}, \quad [G^{++}, D_\nu^{++}] = 0, \\ G^{++} &= h^{++m} \partial_m + h^{(+4)} \mathcal{D}^{--} + h^{(+4)\mu} D_\mu^{--} + h^{++\mu} D_\mu^0 \end{aligned}$$

The transformation of the basic analytic harmonic operator defines transformations of the gravitational superfields

$$\delta \Delta^{++} = -\lambda^{++} \mathcal{D}^0$$

We can also construct the nonanalytic harmonic operator

$$\begin{aligned} \Delta^{--}, \quad \delta \Delta^{--} &= -(\Delta^{--} \lambda^{++}) \Delta^{--} \text{ satisfying the constraints} \\ [(\Delta^{++} - h^{(+4)} \Delta^{--}), \Delta^{--}] &= \mathcal{D}^0 \end{aligned}$$

We define the covariant spinor and vector derivatives

$$\Delta_\alpha^0 = \frac{1}{2} [\Delta^{--}, D_\alpha^{++}], \quad \Delta_\alpha^{--} = [\Delta^{--}, \Delta_\alpha^0], \quad \Delta_a = \frac{i}{2} (\gamma_a)^{\alpha\beta} \{\Delta_\alpha^0, \Delta_\beta^0\}$$

Superconformal compensator hypermultiplet

$$\delta q^+(\zeta) = \Lambda q^+(\zeta), \quad \Lambda = -\frac{1}{2} (\partial_m^A \lambda^m + \partial^{--} \lambda^{++} - \partial_\mu^{--} \lambda^{++\mu} - \partial_\mu^0 \lambda^{0\mu})$$